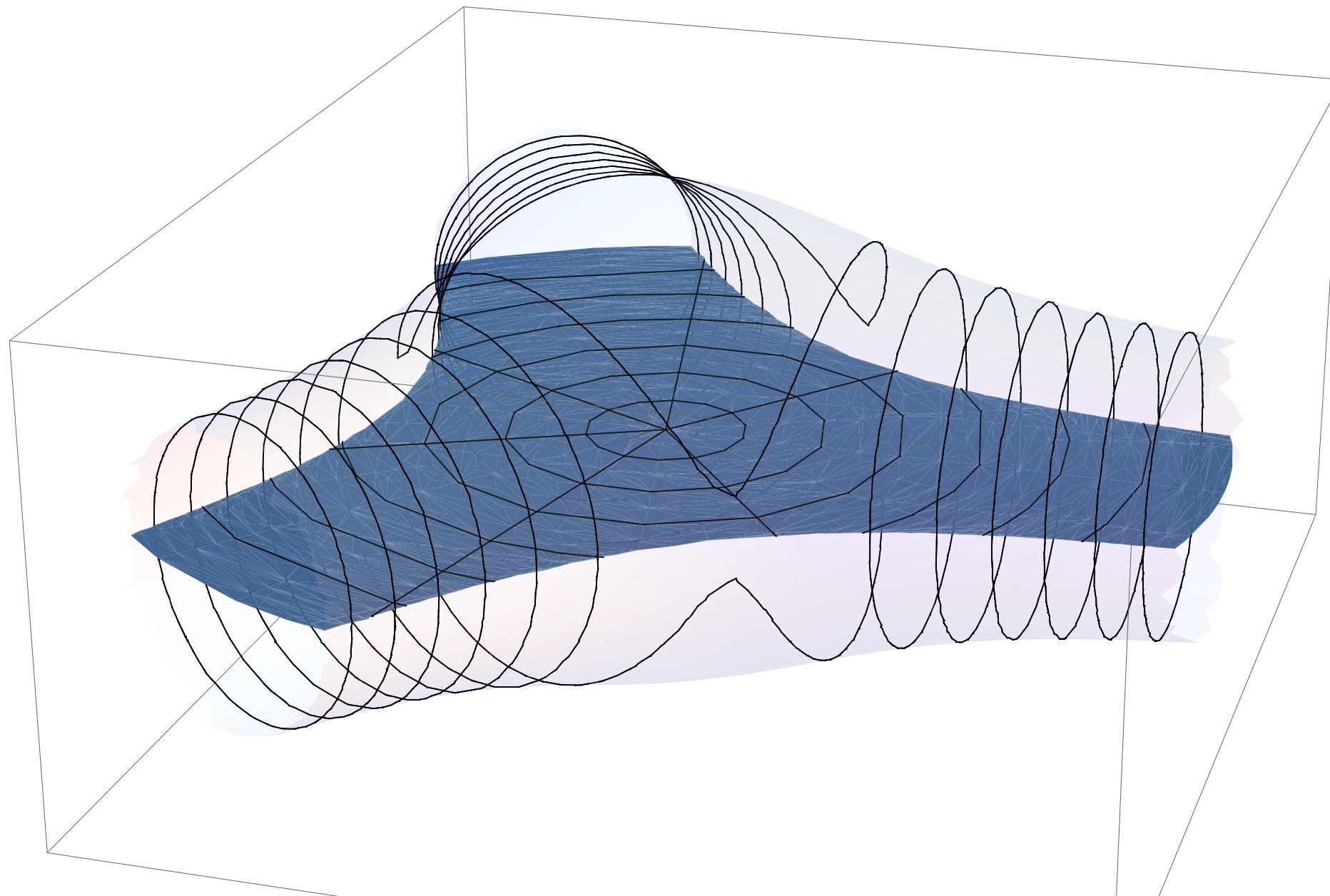


A few recent N-body results

Richard Montgomery
UC Santa Cruz
May 2025 San Luis AMS Conf.



THE NEGATIVE ENERGY N-BODY PROBLEM HAS FINITE DIAMETER

R. Mont. ;

arXiv: 2406.05563; Cel. Mech. and Dyn. Astr: Volume 137, article number 11, (2025)

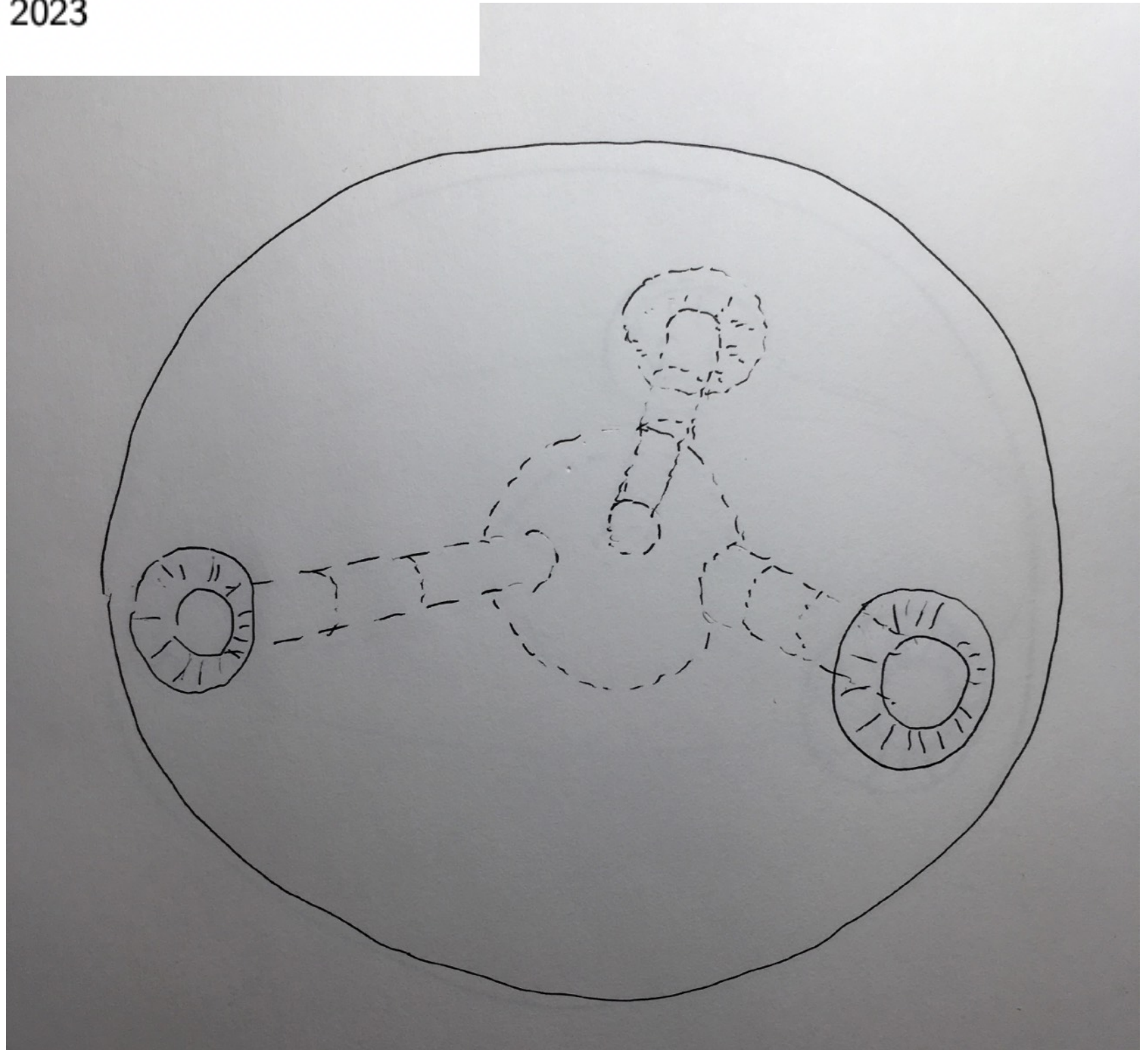
Compactification of the energy surfaces for n bodies

Andreas Knauf* Richard Montgomery[†]

August 15, 2023

[arXiv:2307.03837](https://arxiv.org/abs/2307.03837);

RCD. Oct 2023



Scaling Symmetries, Contact Reduction and Poincaré's dream

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Universidad Nacional Autónoma de México, A. P. 70543, México, DF 04510, Mexico*

Connor Jackman†

CIMAT, A.P. 402, Guanajuato, Gto. 36000, México

David Sloan‡

Department of Physics, Lancaster University, Lancaster UK

A symplectic Hamiltonian system admitting a scaling symmetry can be reduced to an equivalent contact Hamiltonian system in which some physically-irrelevant degree of freedom has been removed.

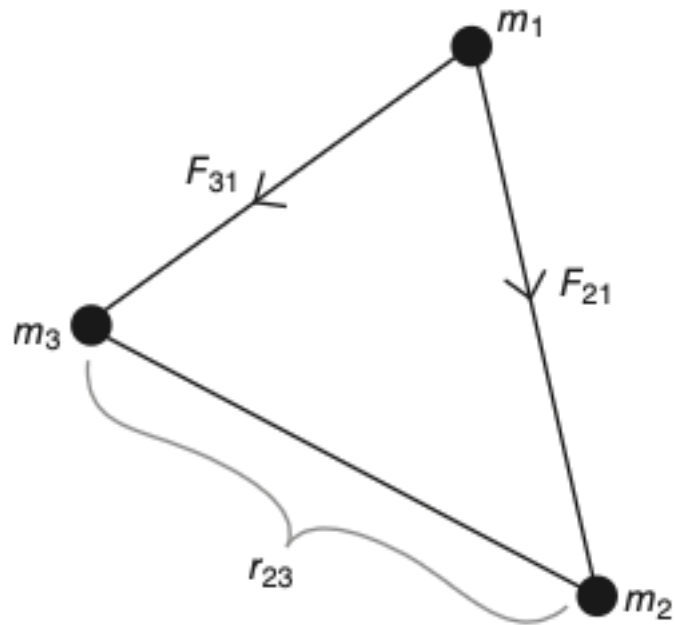
[arXiv:2206.09911](https://arxiv.org/abs/2206.09911)

`Hammers looking for nails'

**These papers represent ``optical instruments’’
which combined, give a picture of the Newtonian
N-body problem as an analytic flow
on a compact stratified analytic space.**

**Will this picture
solve any interesting problems?**

Problem 1: the 'oldest question' of Newton-Herman and parabolic infinity



$$m_1 \ddot{q}_1 = F_{21} + F_{31},$$

$$m_2 \ddot{q}_2 = F_{12} + F_{32},$$

$$m_3 \ddot{q}_3 = F_{23} + F_{13},$$

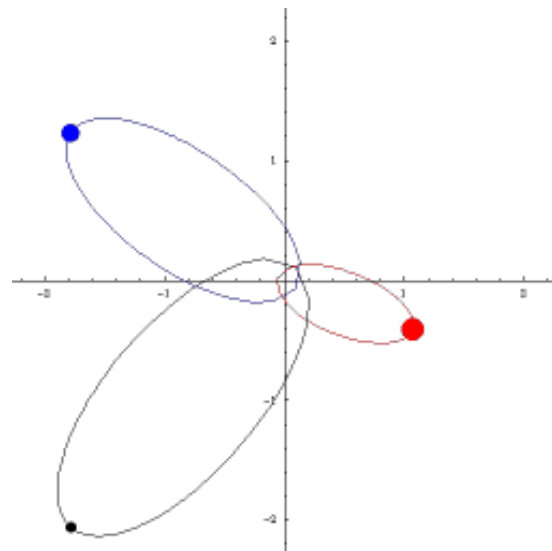
Bounded orbits: all $r_{ab}(t)$ are bounded functions of time

Unbounded orbits: some $r_{ab}(t) \rightarrow \infty$

OLDEST QUESTION: ¿ Do unbounded orbits pass arbitrarily close to any bound orbit in phase space?

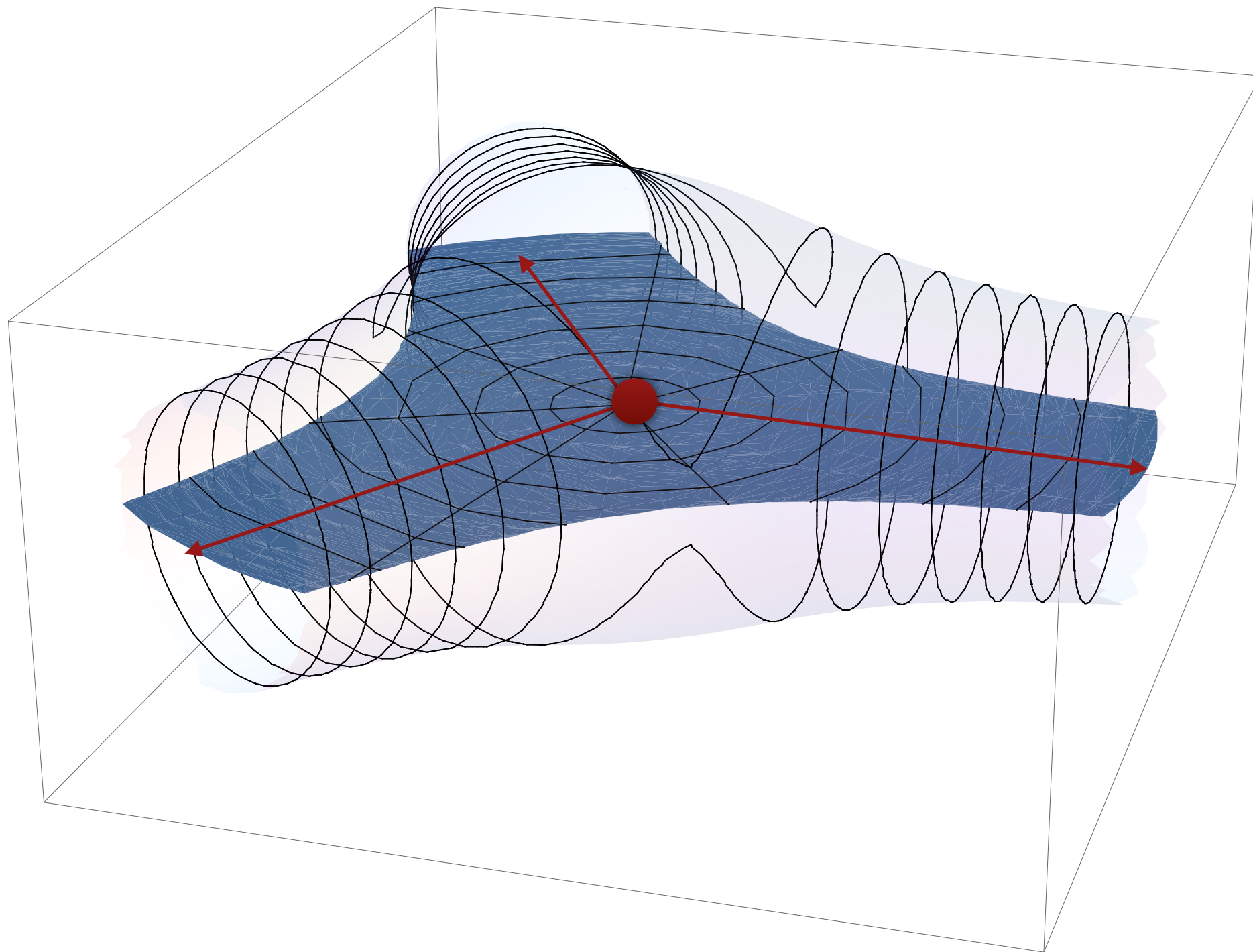
**QUESTION 2: For any given mass ratios,
and ZERO angular momentum do there
exist periodic orbits?**

**If the angular momentum is NON-ZERO:
Yes! Lagrange and Euler's solutions.**

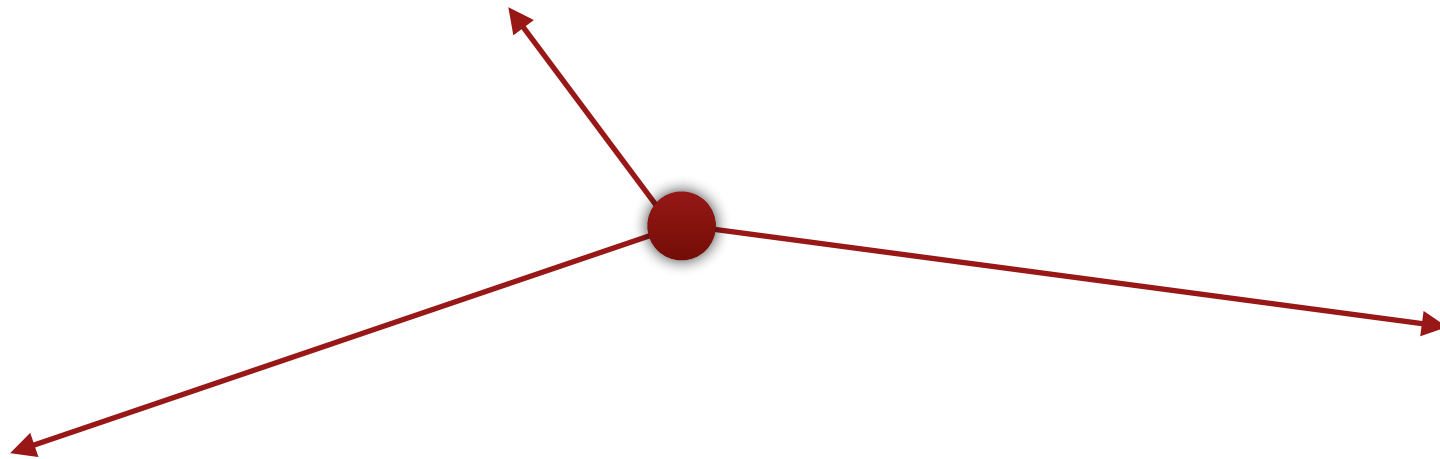


Recall: bounded \implies energy is NEGATIVE.

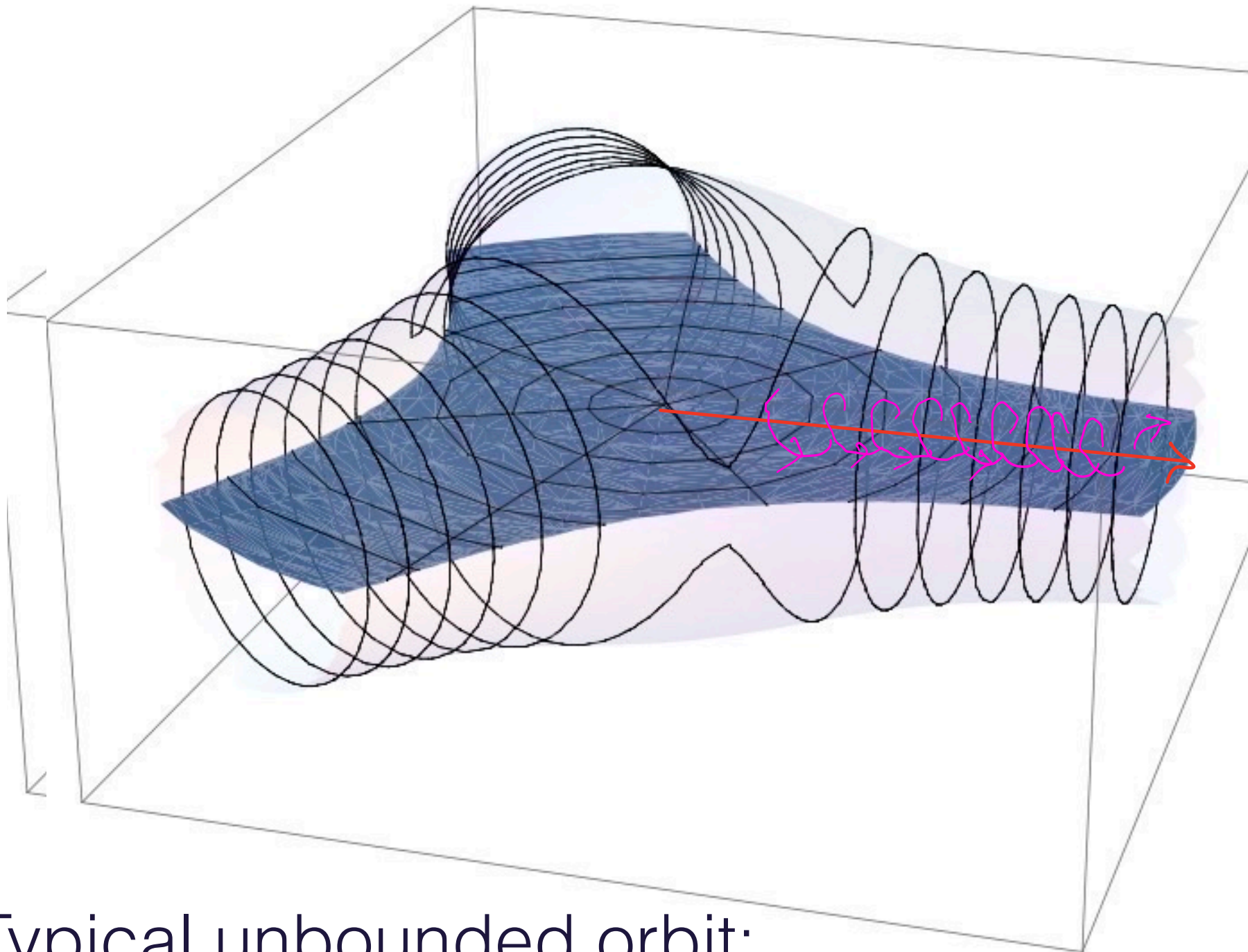
So, for both questions, we are interested in the flow when the energy E is NEGATIVE.



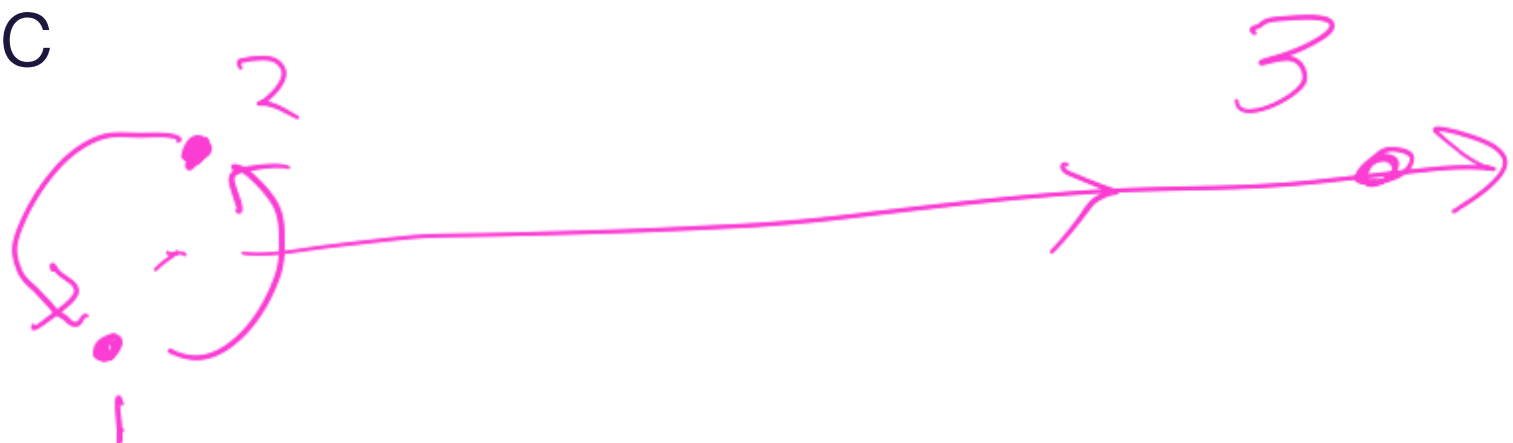
The Hill region at negative energy,
for the planar 3- body problem
projected to shape space



Collision locus=
where potential is negative infinity

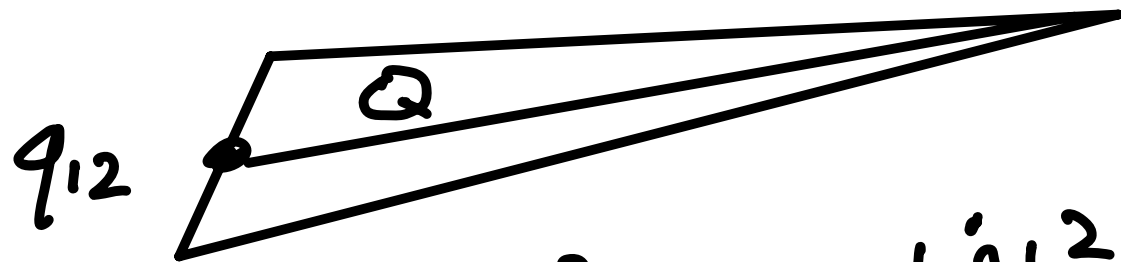
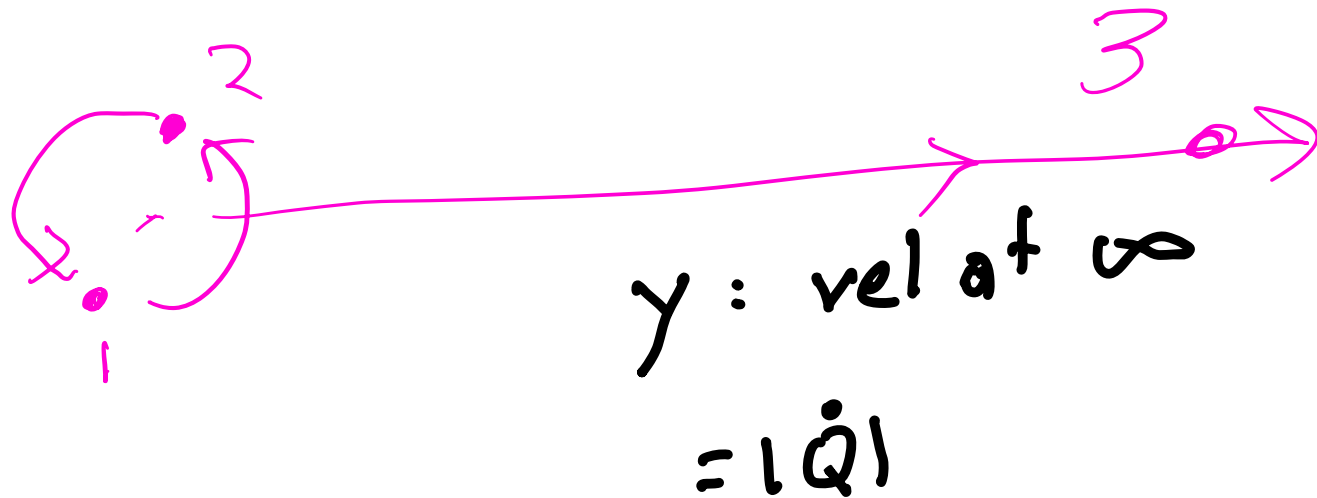


Typical unbounded orbit:
hyperbolic-elliptic

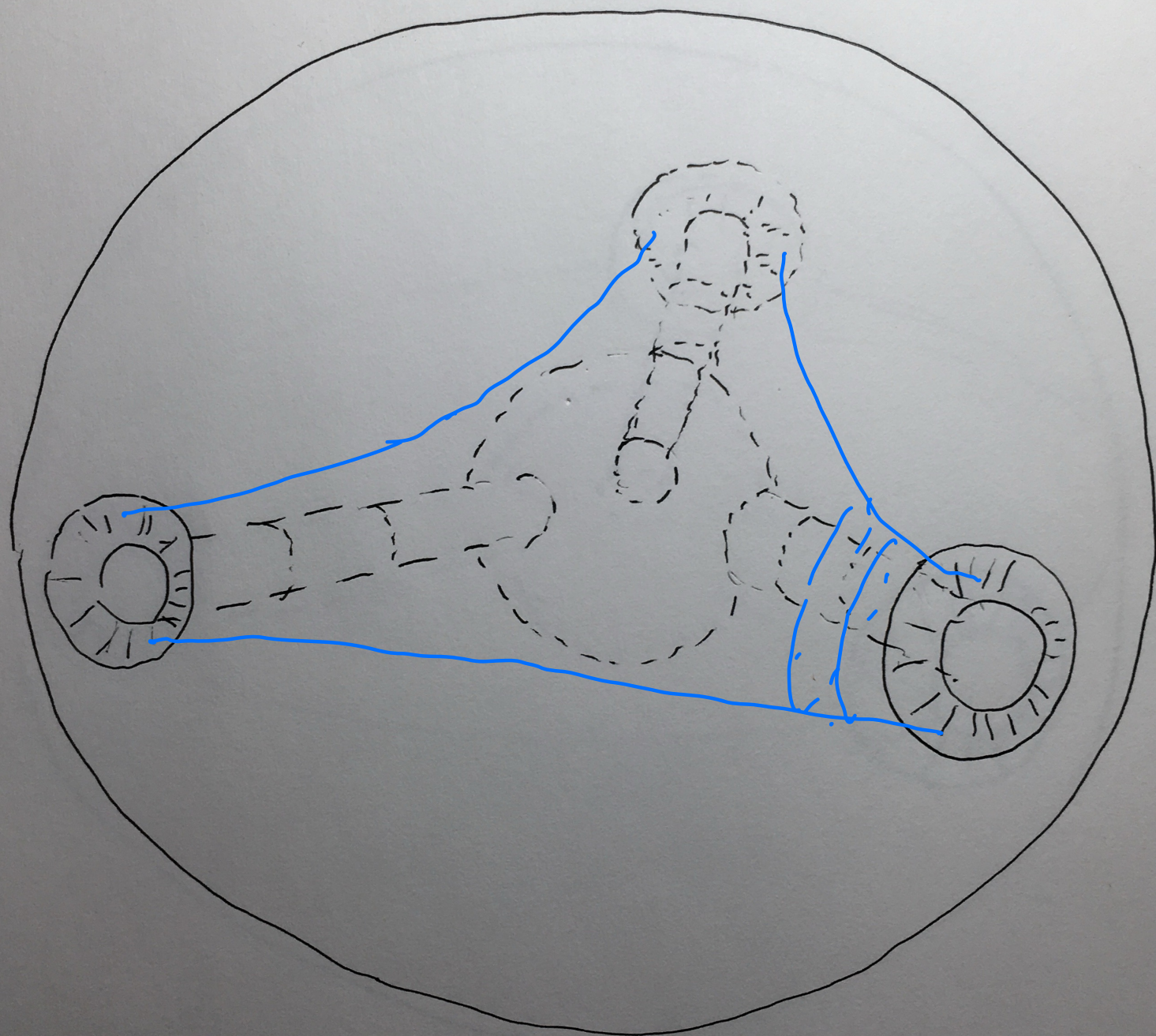


$$(E < 0)$$

'P' : $\rightarrow \gamma = 0$: parabolic-elliptic
 'H' : $\rightarrow \gamma > 0$: hyperbolic-elliptic



$$K = \frac{1}{2} (\mu_1 |\dot{q}_{12}|^2 + \mu_2 |\dot{Q}|^2)$$



After
compactification.

The const. neg. energy hypersurface modulo rigid motions forms a two-sphere bundle over the interior of the blue pair of pants. These spheres degenerate to points on the pants boundary.

Notation and Pictures

Form the non-compact 5-manifold

$$M^5 = M^5(E, J; m_1, m_2, m_3)$$

by fixing the center of mass and linear momentum equal to zero in the planar three-body problem, fixing the energy E and the angular momentum J and forming the quotient of the resulting variety by rotations.

Masses m_1, m_2, m_3 enter as parameters.

Reduced planar 3-body flows live here.

Compactifying M^5 adds boundary strata corresponding to $R \rightarrow \infty$ or $R \rightarrow 0$ or some $r_{ab} \rightarrow 0$.

$$R^2 = I := \sum m_a |q_a|^2 = \sum m_a m_b r_{ab}^2 / \sum m_a$$



Define subsets $B, H, P \subset M^5 = M^5(E, J; m_1, m_2, m_3)$ by:

B = i.c.s of bounded orbits

H = i.c.s for hyperbolic-elliptic orbits

P = i.c.s for parabolic-elliptic orbits.

The answer to the OLDEST Q is 'yes' \iff any of the following equivalent conditions hold

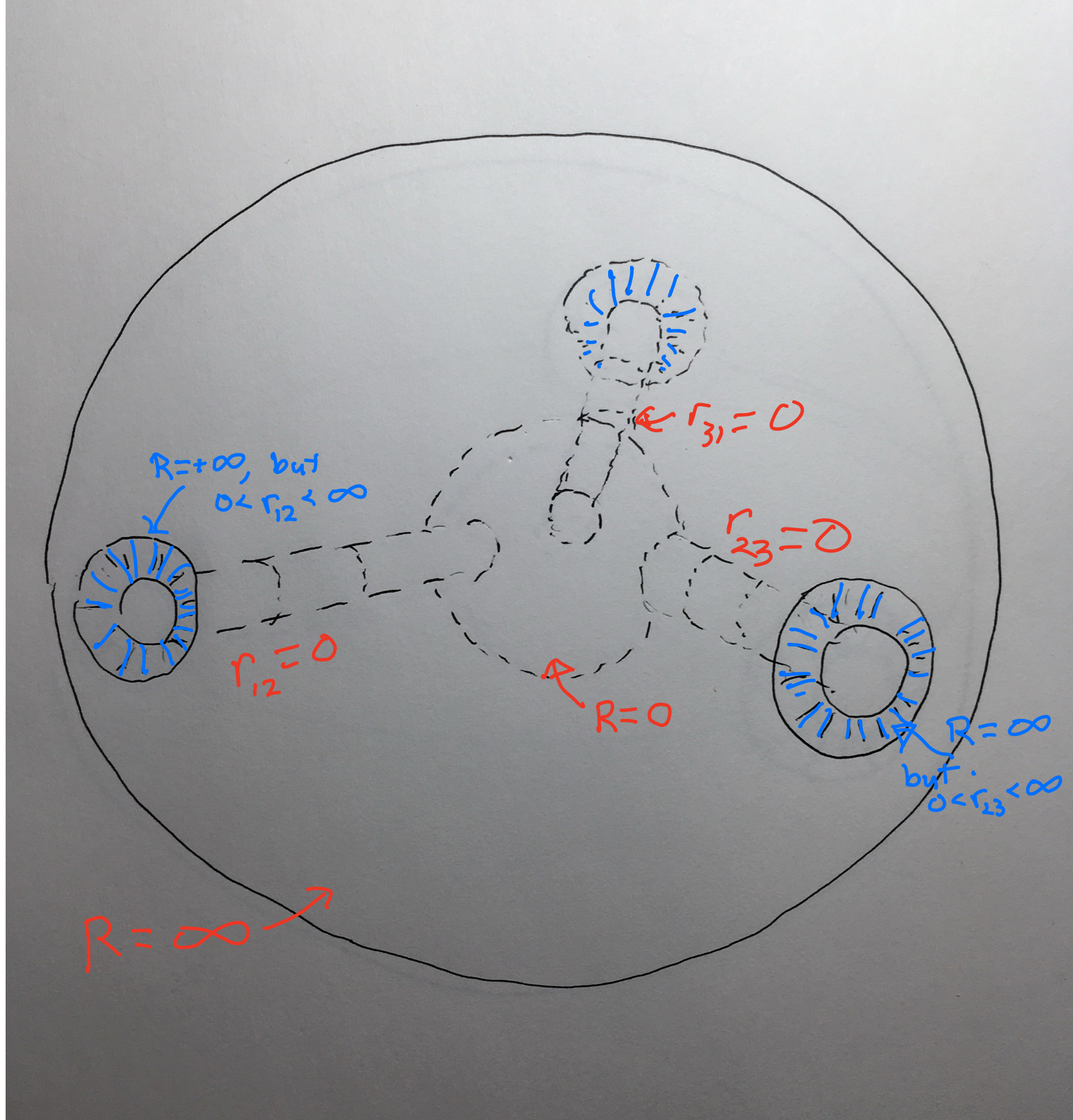
I) $\text{int}(B) = \emptyset$

II) $\bar{H} = M^5$

III) $\bar{P} \supset B$.

What we know: $H \neq \emptyset$. $P \neq \emptyset$. $P \subset \bar{H}$. $B \neq \emptyset$ provided the ang. mom. is nonzero. Often (always??) $\text{meas}(B) > 0$ (KAM).

What we do not know [Q2 !] is $B \neq \emptyset$ when ang. mom. is zero, regardless of mass ratios?



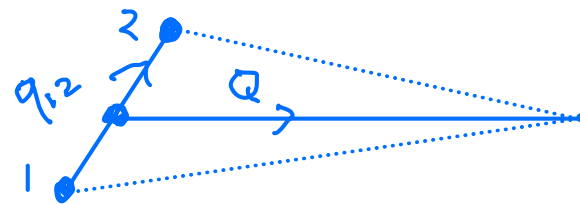
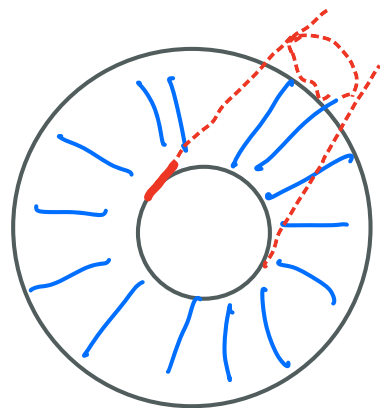
$$R = \sqrt{\sum m_a m_b r_{ab}^2 / \sum m_a}$$

$$= \sqrt{\sum m_a |a|^2}$$

JOURNAL OF DIFFERENTIAL EQUATIONS **52**, 356–377 (1984)

Homoclinic Orbits and Oscillation for the Planar Three-Body Problem

CLARK ROBINSON*

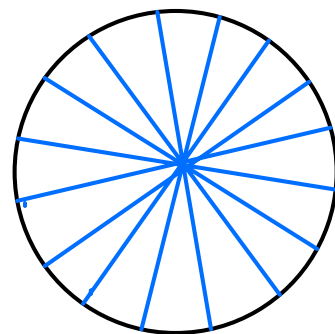


$$R=0 \quad :$$

$$0 < r_{12} < \infty$$

$$E = H_{12}^{\text{Kep}} + H_3^{\text{Kep}} + O\left(\frac{1}{R^2}\right)$$

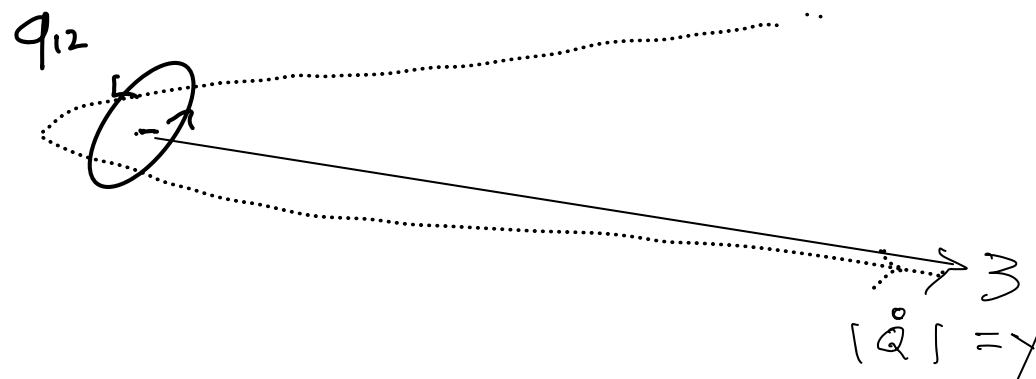
L-C
regularize \Downarrow or Moser



Flow @ ∞ :

$$S^3 \times \mathbb{R}$$

reg.
Kepler γ



The following two theorems give the smoothness of the manifolds $W^s(S^3)$ and $W^u(S^3)$ and the smoothness of the assignment of the asymptotic binary motion. Both are trivially true for the model equations used in [3] where the equations completely decouple for $x^2 + y^2$ small.

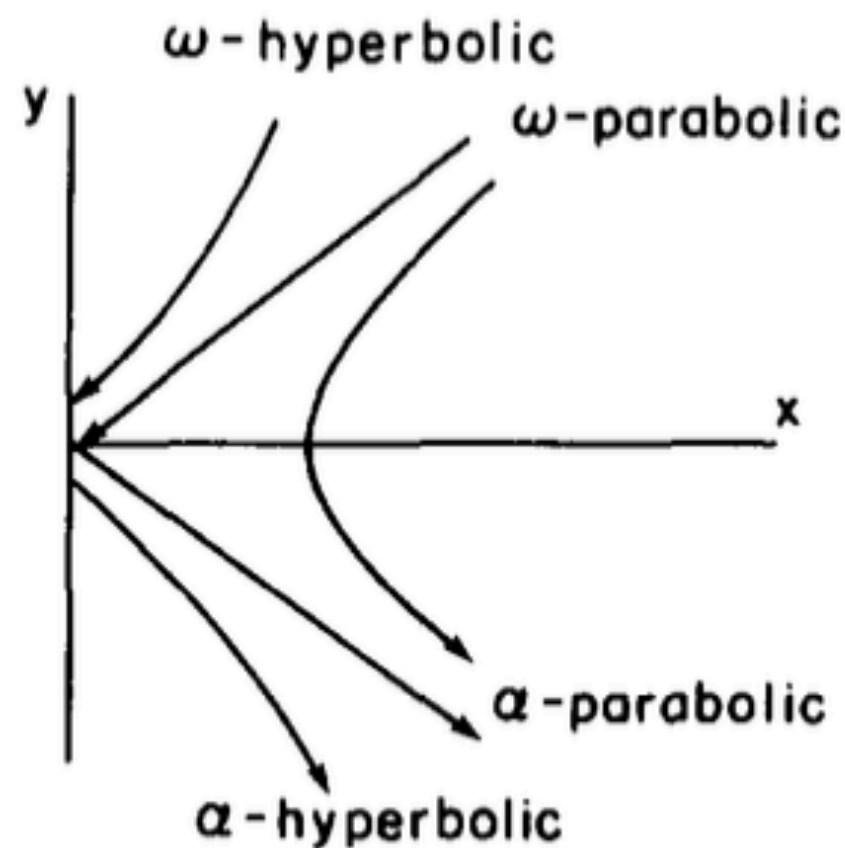
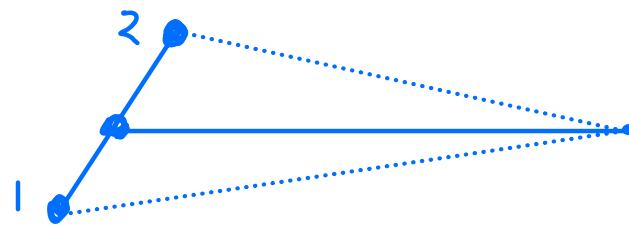
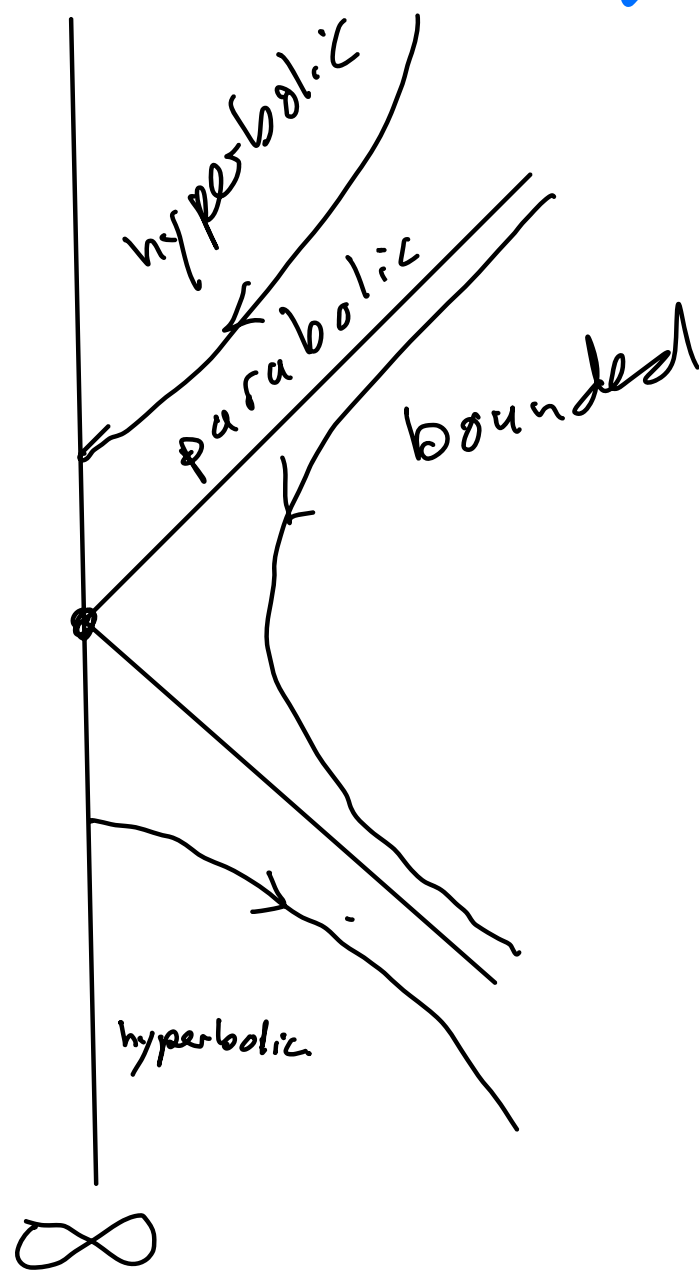


FIGURE 3



- Parabolic ∞ : $S^3 \times \{0\}$
 hopf flow
 all orbits S^1 's

- $W^s(\text{Parab.}) = P = P^4 \hookrightarrow M^5$
 Robinson:
 analytic

- Where does $P^4 = W^s(\text{parab})$ go?

- Does it enter into the interstices
 of my favorite KAM stable orbits?

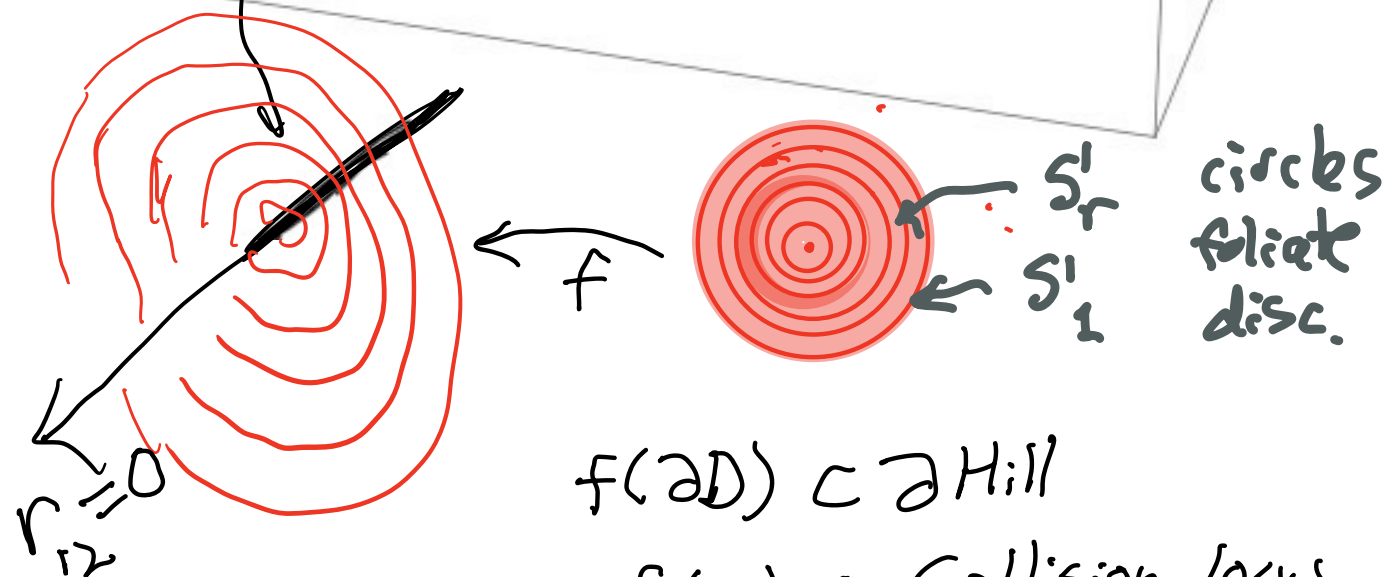
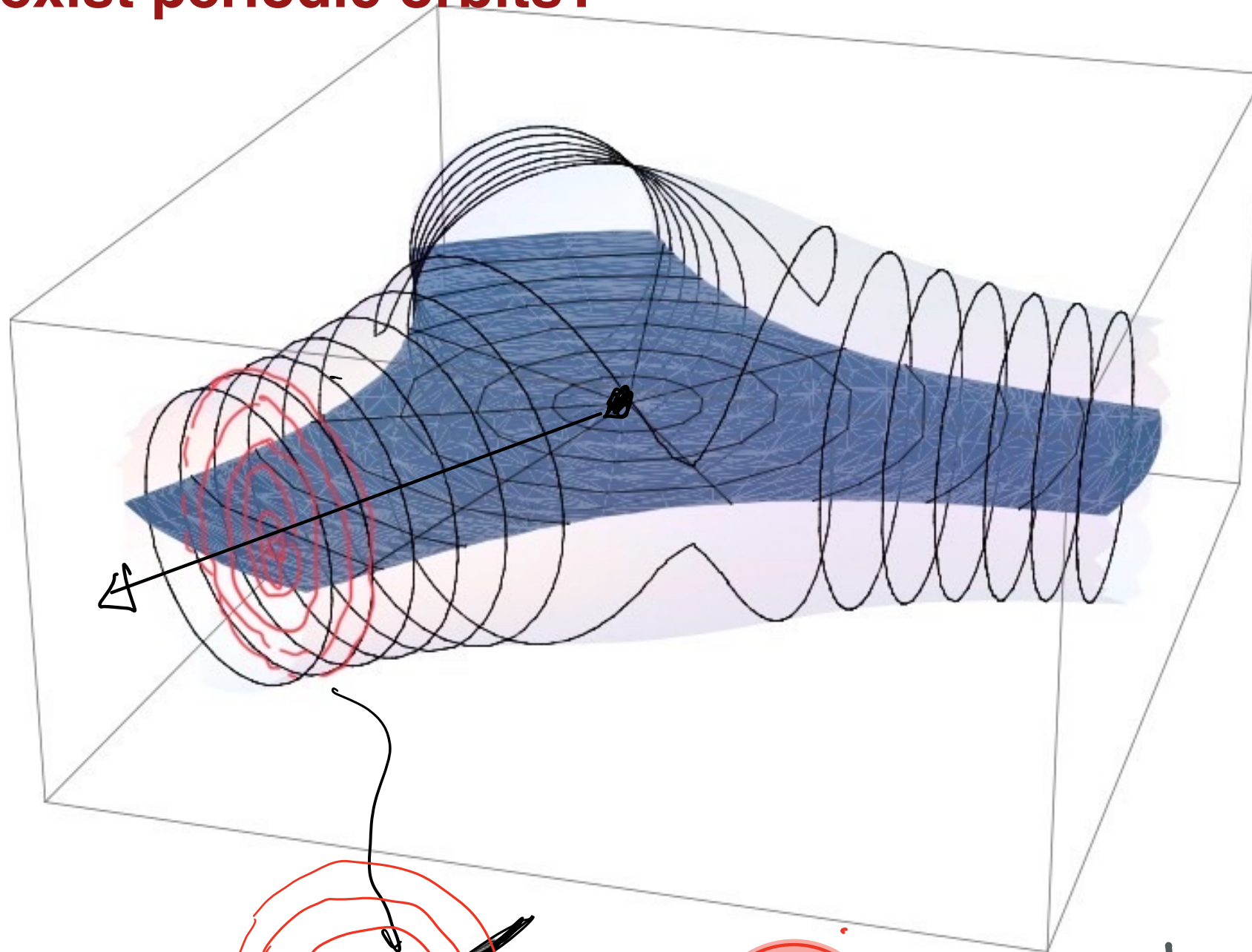
- How so?

- Are variational methods of any use here?

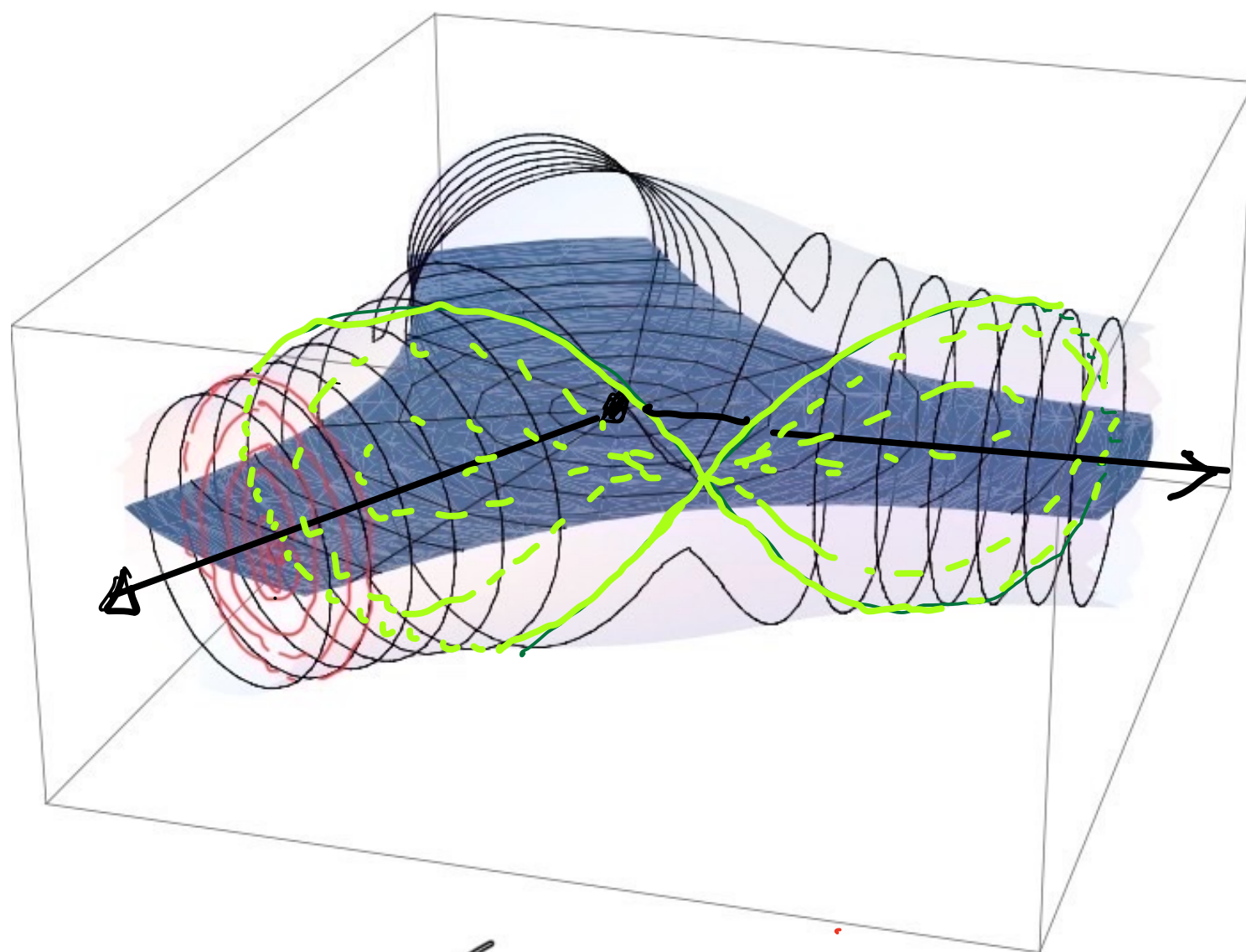
On to Q 2

**QUESTION 2: For any given mass ratios,
and ZERO angular momentum do there
exist periodic orbits?**

**a min-max
attack:**



or :



$$f(\partial D) \subset \partial H$$

$$f(0) = 0$$

But! $f(\partial D)$ realizes
given free homotopy class on ∂H !

Now try min max:

$$\min_f \max_{JM} \ell(f(S_r^+))$$

Some Guesses ...
eg over tight binary class

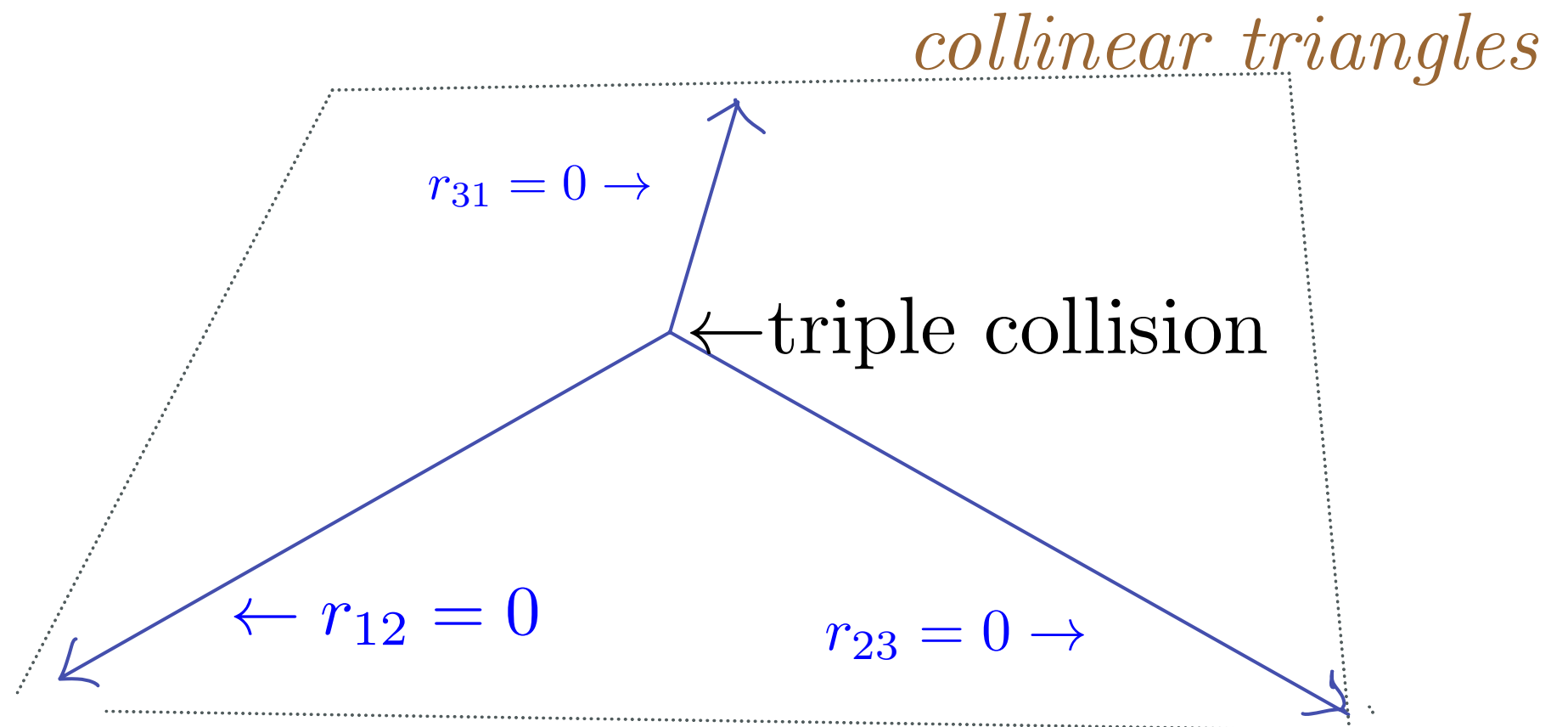
End.

if more time:

- a) details of papers ...
- b) speculations.
- c) audience Qs of course!

SHAPE SPACE.

an \mathbb{R}^3 realizing the quotient
of the planar three-body configuration space
by rigid motions.



Recall: bounded \implies energy is NEGATIVE.